

Ideal Bose gas

Again, there are $g \frac{4\pi p^2 dp}{(2\pi\hbar)^3} V$ states in a spherical layer of width dp

But now there may be >1 particles (bosons) in each state

$$\int g \frac{4\pi p^2 dp}{(2\pi\hbar)^3} \frac{1}{e^{\frac{\epsilon(p)-\mu}{T}} - 1} = \frac{N}{V}$$

Changing variables to $\epsilon(p) = \frac{p^2}{2m}$,

$$\frac{N}{V} = \frac{g m^{\frac{3}{2}}}{2^{\frac{1}{2}} \pi^2 \hbar^3} \int_0^{\infty} \frac{\sqrt{\epsilon} d\epsilon}{e^{\frac{\epsilon-\mu}{T}} - 1}$$

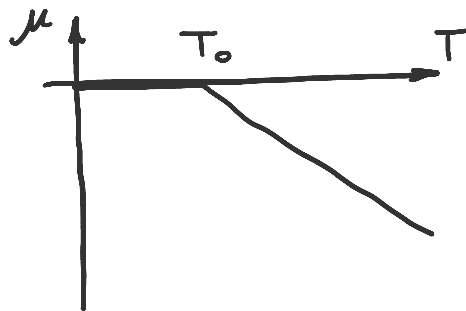
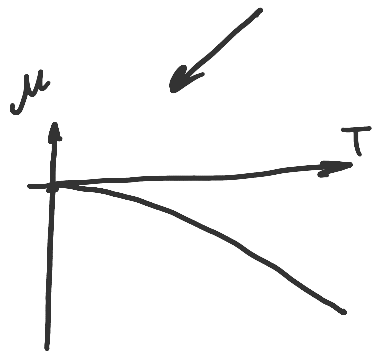
Note an important distinction between Fermi and Bose gases. At $T=0$ a Fermi gas has a finite, while a Bose gas may sit entirely in the state with $E=0$ at $T=0$

(Reminder: $E=0$ is the ground state, and all the bosons will want to be there)

Let's lower the temperature gradually. Will the state with all bosons in the ground state be reached at $T>0$? Assume so.

There are 2 possible scenarios

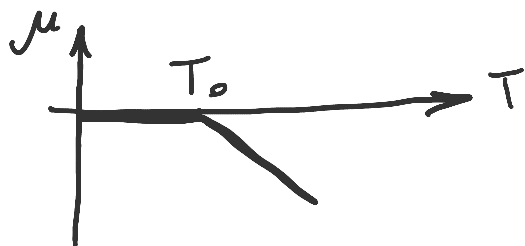
There are 2 possible scenarios



$$\frac{N}{V} = \frac{g (mT)^{\frac{3}{2}}}{2^{\frac{1}{2}} \pi^2 \hbar^3} \underbrace{\int_0^{\infty} \frac{\sqrt{z} dz}{e^z - 1}}_{\frac{\sqrt{\pi}}{2} \zeta(\frac{3}{2})} \quad (z = \frac{\epsilon}{T})$$

$$\frac{N}{V} = \frac{g (mT)^{\frac{3}{2}}}{(2\pi)^{\frac{3}{2}} \hbar^3} \zeta(\frac{3}{2})$$

$$T_0 = \frac{2\pi}{[\zeta(\frac{3}{2})]^{\frac{2}{3}}} \frac{\hbar^2}{m g^{\frac{2}{3}}} \left(\frac{N}{V}\right)^{\frac{2}{3}} \approx \frac{3.31}{g^{\frac{2}{3}}} \frac{\hbar^2}{m} \left(\frac{N}{V}\right)^{\frac{2}{3}}$$



What about the condition of having $\mu < 0$ for all temperatures in a Bose gas?

One has to be more careful with μ . There is some uncertainty there.

1. Is $E = 0$. There is some uncertainty there.

2. Is μ how many

the level $E=0$. There is some uncertainty we cannot tell from this formula how many particles there are on the level with $E=0$.

But let us count the particles with $E > 0$

$$\frac{g V (mT)^{\frac{3}{2}}}{2^{\frac{1}{2}} \pi^2 \hbar^3} \int_0^{\infty} \frac{\sqrt{z} dz}{e^z - 1} = N \cdot \left(\frac{T}{T_0}\right)^{\frac{3}{2}}$$

(for $\mu=0$)

$$N_{E=0} = N \left[1 - \left(\frac{T}{T_0}\right)^{\frac{3}{2}} \right]$$

BEC is often related to superconductivity and superfluidity (e.g. He^4).

In 1995 observed explicitly by
E. Cornell and C. Wieman

// The most famous picture

Let's compute the energy at $T < T_0$.

$$E = \frac{g V (mT)^{\frac{3}{2}}}{2^{\frac{1}{2}} \pi^2 \hbar^3} \int_0^{\infty} \frac{\varepsilon^{\frac{3}{2}} d\varepsilon}{e^{\frac{\varepsilon}{T}} - 1} =$$

$$= \frac{g V (mT)^{\frac{3}{2}} T}{2^{\frac{1}{2}} \pi^2 \hbar^3} \int_0^{\infty} \frac{z^{\frac{3}{2}} dz}{e^z - 1} = \frac{3g m^{\frac{3}{2}} T^{\frac{5}{2}} V}{2^{\frac{5}{2}} \pi^{\frac{3}{2}} \hbar^3} \zeta\left(\frac{5}{2}\right)$$

≈ 1.51

$$2^{-1/2} \pi^{3/2} \frac{3\sqrt{\pi}}{4} \zeta\left(\frac{5}{2}\right)$$

The specific heat is $C_V = \left(\frac{dE}{dT}\right)_V = \frac{5}{2} \frac{E}{T}$

$$\left. \begin{array}{l} \text{Thus, } C_V \propto T^{\frac{3}{2}} \\ C_V = T \left(\frac{dS}{dT}\right)_V \end{array} \right\} \rightarrow S = \frac{2}{3} \frac{5}{2} \frac{E}{T} = \frac{5}{3} E$$

$$F = E - TS = -\frac{2}{3} E$$

(Note that the number of particles is conserved)

$$\text{So, } F = - \frac{g m^{\frac{3}{2}} T^{\frac{5}{2}} V}{2^{\frac{5}{2}} \pi^{\frac{3}{2}} \hbar^3} \zeta\left(\frac{5}{2}\right)$$

The pressure

$$P = - \left(\frac{\partial F}{\partial V}\right)_T = \frac{g m^{\frac{3}{2}} \zeta\left(\frac{5}{2}\right)}{2^{\frac{5}{2}} \pi^{\frac{3}{2}} \hbar^3} T^{\frac{5}{2}}$$